Exact propagator and eigenfunctions for multistable models with arbitrarily prescribed N lowest eigenvalues

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# Exact propagator and eigenfunctions for multistable models with arbitrarily prescribed $\boldsymbol{N}$ lowest eigenvalues 

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#### Abstract

We present a method to construct potentials for Schrödinger equations with some prescribed features, for which all eigenfunctions and the time-dependent propagator can be explicitly calculated. The prescribed features can be formulated by choosing arbitrarily the $N$ lowest eigenvalues. Alternatively one can prescribe some qualitative behaviour for the potential, like the number and relative depths of wells and barriers. The results can also be applied to the construction of Fokker-Planck models with prescribed properties and explicitly calculable transition probability density. The method is based on ideas of supersymmetric quantum mechanics and the theory of solitons, that can be traced back to the work of Darboux and Crum.


## 1. Introduction

Many properties of quantum systems, such as those at low temperatures, depend essentially only on the shape of the potential at low energies, or on the lowest eigenvalues and eigenfunctions. In general it is not possible, even in one-dimensional models, to determine the propagator and eigenfunctions for a given problem. It is thus of interest to be able to construct potentials that have a prescribed set of $N$ lowest eigenvalues, for which these quantities can be calculated exactly. The problem can also be formulated in terms of the construction of a potential having some prescribed qualitative features, such as the number and depth of wells, or even some approximate quantitative behaviour at low energies.

The present approach starts with a potential $V_{0}$ for which the corresponding propagator and eigenfunctions are explicitly known. $V_{0}$ is then modified in such a way as to add $N$ arbitrary eigenvalues at the bottom of the spectrum. This step involves the knowledge of the solutions of the original Schrödinger equation at energies that are smaller than the bottom of the spectrum. The propagator and eigenfunctions for the new potential $V_{N}$ can be expressed in terms of the original ones, involving only derivatives and integrals. Figure 1 shows some examples of potentials $V_{N}$ constructed with this method starting from $V_{0}=0$.

In a closely related context, the method allows one to construct multistable FokkerPlanck models, for which the time-dependent transition probabilities can be calculated explicitly.

The basic ideas of this method can be traced back to the work of Darboux [1] and Crum [2]. These ideas have found applications in supersymmetric quantum mechanics [3] and in the study of solitons and inverse scattering theory [4,5]. Some special cases of this method have been treated in the iiterature: eigenfunctions have been calculated


Figure 1. (a) Potentials $V_{N}$ constructed by adding to $V_{0}=0$ the following eigenvalues $\lambda_{i}$ : (A) $-1,-1.66$; (B) $-1,-1.66,-2.637$; (C) $-1,-1.66,-2.766,-3.269$. (The curves are shifted in the $V_{N}$ direction for clarity.) (b) Potential $V_{N}$ constructed by adding to $V_{0}=0$ the eigenvalues $\lambda_{i}=-\mu_{i}^{2}$, for $\mu_{i}=1,1.2,1.4,1.6,1.8,2.0,2.2$.
in [6-9], and time-dependent solutions and propagators were obtained in [10, 11]. In [12,13] expressions for transition probabilities in bistable Fokker-Planck models were obtained for particular initial conditions.

## 2. Darboux-Crum transformation

We consider a Schrödinger operator (in the Hilbert space $L_{2}(R, \mathrm{~d} x)$ )

$$
\begin{equation*}
H_{-}=-\frac{\mathrm{d}^{2}}{\mathrm{~d} x^{2}}+V_{-}(x) \tag{2.1}
\end{equation*}
$$

that has a ground state $\varphi_{0}$ with energy $E_{0}$, which we write as

$$
\begin{equation*}
\varphi_{0}=\mathrm{e}^{-W} . \tag{2.2}
\end{equation*}
$$

Then $H_{-}$admits the following representation (with primes denoting derivatives)

$$
\begin{equation*}
\left(H_{-}-E_{0}\right)=A^{\dagger} A \tag{2.3}
\end{equation*}
$$

with

$$
\begin{align*}
& A=\frac{\mathrm{d}}{\mathrm{~d} x}+W^{\prime} \equiv \varphi_{0} \frac{\mathrm{~d}}{\mathrm{~d} x} \varphi_{0}^{-1} \\
& A^{+}=-\frac{\mathrm{d}}{\mathrm{~d} x}+W^{\prime} \equiv-\varphi_{0}^{-1} \frac{\mathrm{~d}}{\mathrm{~d} x} \varphi_{0} \tag{2.4}
\end{align*}
$$

We define a new operator

$$
\begin{equation*}
H_{+} \doteq A A^{\dagger}+E_{0}=-\frac{\mathrm{d}^{2}}{\mathrm{~d} x^{2}}+V_{+} \tag{2.5}
\end{equation*}
$$

where

$$
\begin{equation*}
V_{+}=V_{-}+2 W^{\prime \prime} \equiv-V_{-}+2 E_{0}+2 W^{\prime 2} \tag{2.6}
\end{equation*}
$$

Remark that $V_{ \pm}$can also be written as

$$
\begin{equation*}
V_{-}=W^{\prime 2}-W^{\prime \prime}+E_{0} \quad V_{+}=W^{\prime 2}+W^{\prime \prime}+E_{0} . \tag{2.7}
\end{equation*}
$$

$H_{+}$is called the supersymmetric partner of $H_{-}$. Their spectral properties are related as follows.
(1) The spectrum of $H_{+}$is equal to that of $H_{-}$but without the eigenvalue $E_{0}$.
(2) If $f(x)$ satisfies the equation

$$
\begin{equation*}
H_{-} f=\Gamma f \tag{2.8}
\end{equation*}
$$

where $\Gamma$ can be
either a real number $\lambda$ or the operator $\sigma \frac{\mathrm{d}}{\mathrm{d} t} \quad$ with $\sigma=\left\{\begin{array}{l}\mathrm{i} \\ -1\end{array}\right.$
then $g \doteq A f$ satisfies

$$
\begin{equation*}
H_{+} g=\Gamma g . \tag{2.10}
\end{equation*}
$$

Conversely, if $g$ satisfies (2.10) then $h \neq A^{+} g$ satisfies (2.8).
The two values of $\sigma=\mathrm{i},-1$ correspond to the Schrödinger and the diffusion ( $\sim$ Fokker-Planck) equations respectively.

Property (1) is proven in [4]. The proof can be sketched as follows. First one shows that $E_{0}$ is not an eigenvalue of $H_{+}$: by (2.4) the only candidate for eigenfunction is $\varphi_{0}^{-1}$, which is not normalisable. Then one considers the polar decomposition

$$
\begin{equation*}
A=U|A| \quad A^{\dagger}=|A| U^{\dagger} \tag{2.11}
\end{equation*}
$$

which gives

$$
\begin{align*}
& H_{-}-E_{0}=A^{\dagger} A=|A|^{2}  \tag{2.12}\\
& H_{+}-E_{0}=A A^{\dagger}=U|A|^{2} U^{\dagger}=U\left(H_{-}-E_{0}\right) U^{\dagger} \tag{2.13}
\end{align*}
$$

where $|A| \doteq\left(A^{\dagger} A\right)^{1 / 2}$, and $U$ is isometric ( $U^{\dagger}=U^{-1}$ ) and can be written as

$$
\begin{equation*}
U=A\left(H_{-}-E_{0}\right)^{-1 / 2} \tag{2.14}
\end{equation*}
$$

As (2.14) indicates, $U$ is defined only in $K_{0}^{\perp}$, the orthogonal complement of the eigenspace of $H_{-}$corresponding to $E_{0}$. Moreover, $U$ maps $K_{0}^{+}$onto the whole space $L_{2}$ :

$$
\begin{align*}
& U: K_{0}^{\perp} \rightarrow L_{2} \\
& U^{\dagger}: L_{2} \rightarrow K_{0}^{\perp} . \tag{2.15}
\end{align*}
$$

Thus, (2.13) states the unitary equivalence between $H_{+}$and the restriction of $H_{-}$to the subspace $K_{0}^{\perp}$, which implies property (1).

Property (2) is verified by insertion.
This allows us to relate the normalised eigenfunctions $\varphi_{(+), n}$ and $\varphi_{(-), n}($ for $n \geqslant 1)$ corresponding to the common eigenvalue $E_{n}$ of $H_{-}$and $H_{+}$:

$$
\begin{align*}
& \boldsymbol{\varphi}_{(-), n}=\left(E_{n}-E_{0}\right)^{-1 / 2} A^{\dagger} \boldsymbol{\varphi}_{(+), n}  \tag{2.16}\\
& \boldsymbol{\varphi}_{(+), n}=\left(E_{n}-E_{0}\right)^{-1 / 2} A \boldsymbol{\varphi}_{(-), n} . \tag{2.17}
\end{align*}
$$

Property (2) allows us to express the time-dependent solutions associated with $H_{-}$in terms of time-dependent solutions associated with $H_{+}$, and vice versa.

## 3. Adding one eigenvalue

We start with a Schrödinger operator

$$
\begin{equation*}
H_{0}=-\frac{\mathrm{d}^{2}}{\mathrm{~d} x^{2}}+V_{0}(x) \tag{3.1}
\end{equation*}
$$

with a potential that grows asymptotically at $\pm \infty$ as

$$
\begin{equation*}
V_{0} \sim c_{ \pm} x^{\eta_{ \pm}} \quad c_{ \pm}, \eta_{ \pm} \geqslant 0 \tag{3.2}
\end{equation*}
$$

We also require that the asymptotic behaviour of its first $N$ derivatives is given by differentiating both sides of (3.2). This includes potentials with a purely discrete spectrum as well as, for example, the free particle. We will denote by $\lambda_{0}$ the lowest point of the spectrum.

We will construct a new Hamiltonian $H_{1}$ that has the same spectrum as $H_{0}$ but with one supplementary eigenvalue $\lambda_{1}<\lambda_{0}$. The idea is to identify ( $H_{0}-\lambda_{1}$ ) with the ( $H_{+}-E_{0}$ ) of the last section; the corresponding $H_{-}$is then identified as the new $H_{1}$. We want to represent $H_{0}$ as

$$
\begin{equation*}
\left(H_{0}-\lambda_{1}\right)=A_{1} A_{1}^{\dagger} \tag{3.3}
\end{equation*}
$$

with

$$
\begin{equation*}
A_{1}=\varphi_{(1), 0} \frac{\mathrm{~d}}{\mathrm{~d} x}\left(\varphi_{(1), 0}\right)^{-1}=\frac{\mathrm{d}}{\mathrm{~d} x}+W_{1}^{\prime} . \tag{3.4}
\end{equation*}
$$

Then

$$
\begin{equation*}
H_{1} \equiv-\frac{\mathrm{d}^{2}}{\mathrm{~d} x^{2}}+V_{1}(x) \doteq A_{1}^{\dagger} A_{1}+\lambda_{1} \tag{3.5}
\end{equation*}
$$

will have the desired properties.
Equation (3.3) together with (3.4) is equivalent to

$$
\begin{align*}
V_{0}-\lambda_{1} & =\left(W_{1}^{\prime}\right)^{2}+W_{1}^{\prime \prime} \\
& \equiv \varphi_{(1), 0} \frac{\mathrm{~d}^{2}}{\mathrm{~d} x^{2}} \varphi_{(1), 0}^{-1} \tag{3.6}
\end{align*}
$$

or further

$$
\begin{equation*}
\left(-\frac{\mathrm{d}^{2}}{\mathrm{~d} x^{2}}+V_{0}\right) \varphi_{(1), 0}^{-1}=\lambda_{1} \varphi_{(1), 0}^{-1} . \tag{3.7}
\end{equation*}
$$

Thus, all that is needed in order to add one eigenvalue is to find a positive solution of the original Schrödinger equation (corresponding to the Hamiltonian (3.1)) for a $\lambda_{1}<\lambda_{0}$, such that its inverse is normalisable. As we will see, for each $\lambda_{1}$ there is a one-parameter family of such solutions. The new potential is given by

$$
\begin{align*}
V_{1} & =V_{0}-2 W_{1}^{\prime \prime} \\
& =-V_{0}+2 \lambda_{1}+2\left(\varphi_{(1), 0} \frac{\mathrm{~d}}{\mathrm{~d} x} \varphi_{(1), 0}^{-1}\right)^{2} \tag{3.8}
\end{align*}
$$

## 4. Asymptotic properties

In this section we state some results on the existence of solutions of (3.7) that satisfy the required conditions, and characterise them completely. More details and sketches of the proofs are given in the appendix.
(1) There are two linearly independent positive solutions of (3.7) that are uniquely defined by the asymptotic conditions

$$
\begin{align*}
& h_{1}(x) \xrightarrow[x \rightarrow-\infty]{\longrightarrow \longrightarrow \longrightarrow \longrightarrow}\left(V_{0}-\lambda_{1}\right)^{-1 / 4} \exp \left(\int^{x} \mathrm{~d} y\left(V_{0}-\lambda_{1}\right)^{1 / 2}\right) \rightarrow 0  \tag{4.1}\\
& h_{3}(x) \underset{x \rightarrow+\infty}{\longrightarrow \times m \longrightarrow}\left(V_{0}-\lambda_{1}\right)^{-1 / 4} \exp \left(-\int^{x} \mathrm{~d} y\left(V_{0}-\lambda_{1}\right)^{1 / 2}\right) \rightarrow 0 \tag{4.2}
\end{align*}
$$

$h_{1}$ and $h_{3}$ go exponentially to $\infty$ in the opposite limits (see appendix, equations (A6) and (A7)).
(2) With these two functions we can construct a one-parameter family of solutions

$$
\begin{equation*}
g_{0}\left(x ; \lambda_{1}, \alpha_{1}\right) \doteq \alpha_{1} h_{1}(x)+h_{3}(x) \tag{4.3}
\end{equation*}
$$

that, for $\alpha_{1}>0$, are everywhere positive and whose inverse is square integrable. We can thus use $g_{0}$ for the addition of an eigenvalue. This family, together with $h_{1}$ and $h_{3}$, are all the positive solutions of (3.7) (up to multiplication with a constant), because $h_{1}$ and $h_{3}$ are linearly independent, and a negative $\alpha_{1}$ would lead to negative values for $x \rightarrow \infty$.

Remark that if $V_{0}$ is symmetric and we set $\alpha_{1}=1$ then $g_{0}$ is the (unique) even solution, whereas $\alpha_{1}=-1$ gives the odd solution.
(3) The new potential $V_{1}$ has the same asymptotic behaviour (3.2) as $V_{0}$. This is a consequence of (3.8) and property ( $k$ ) of the appendix.

## 5. Adding two eigenvalues

One could now iterate the procedure of $\S 3$ and add eigenvalues $\lambda_{1}, \lambda_{2}, \ldots, \lambda_{N}$ successively. This involves the determination of a positive solution of a new Schrödinger equation at each step. Since the potentials become more and more complicated this may look like a hopeless task. However, it turns out that it is sufficient to know the solutions $h_{1}(x ; \lambda)$ and $h_{3}(x ; \lambda)$ of the original $H_{0}$ for $\lambda=\lambda_{1}, \lambda_{2}, \ldots, \lambda_{N}$, and that one can construct $H_{N}$ in one step.

In order to explain the procedure, we consider first the addition of a second eigenvalue $\lambda_{2}$. Let $g_{0}\left(x ; \lambda_{1}, \alpha_{1}\right)$ be a positive solution for $\lambda=\lambda_{1}$ of

$$
\begin{equation*}
H_{0} f=\lambda f \tag{5.1}
\end{equation*}
$$

as defined in (4.3). Define

$$
\begin{equation*}
V_{1}=V_{0}-2 \frac{\mathrm{~d}^{2}}{\mathrm{~d} x^{2}} \ln g_{0} . \tag{5.2}
\end{equation*}
$$

We know from $\S 2$ that, if $f_{0}(x ; \lambda)$ is a solution of (5.1) for an arbitrary $\lambda$, then a solution of

$$
\begin{equation*}
H_{1} f=\lambda f \tag{5.3}
\end{equation*}
$$

is given by

$$
\begin{align*}
f_{1}(x ; \lambda) & \doteq c A_{1}^{\dagger} f_{0}(x ; \lambda)  \tag{5.4}\\
& \equiv-c g_{0}^{-1}\left(\lambda_{1}, \alpha_{1}\right) \operatorname{det}\left(\begin{array}{ll}
g_{0}\left(\lambda_{1}, \alpha_{1}\right) & f_{0}(\lambda) \\
g_{0}^{\prime}\left(\lambda_{1}, \alpha_{1}\right) & f_{0}^{\prime}(\lambda)
\end{array}\right)  \tag{5.5}\\
& \equiv-c \frac{\Omega_{2}\left(g_{0}\left(\lambda_{1}, \alpha_{1}\right), f_{0}(\lambda)\right)}{\Omega_{1}\left(g_{0}\left(\lambda_{1}, \alpha_{1}\right)\right)} \tag{5.6}
\end{align*}
$$

where $c$ is a constant and

$$
\begin{equation*}
A_{1}^{\dagger}=-g_{0}\left(\lambda_{1}, \alpha_{1}\right) \frac{\mathrm{d}}{\mathrm{~d} x} g_{0}^{-1}\left(\lambda_{1}, \alpha_{1}\right) \tag{5.7}
\end{equation*}
$$

In (5.6) we have introduced the notation

$$
\Omega_{n}\left(f_{1}, f_{2}, \ldots, f_{n}\right) \doteq \operatorname{det}\left(\begin{array}{ccc}
f_{1} & \cdots & f_{n}  \tag{5.8}\\
f_{1}^{\prime} & \cdots & f_{n}^{\prime} \\
\vdots & \ddots & \vdots \\
f_{1}^{(n-1)} & & f_{n}^{(n-1)}
\end{array}\right)
$$

and $\Omega_{0} \doteq 1$. In fact, all the solutions can be obtained this way. The general positive solution $g_{1}$ is given by

$$
\begin{equation*}
g_{1}=\frac{\Omega_{2}\left(g_{0}\left(\lambda_{1}, \alpha_{1}\right), g_{0}\left(\lambda_{2},-\alpha_{2}\right)\right)}{\Omega_{1}\left(g_{0}\left(\lambda_{1}, \alpha_{1}\right)\right)} \tag{5.9}
\end{equation*}
$$

where $\alpha_{2}>0$ is an arbitrary constant. Notice that the sign in $-\alpha_{2}$ has the effect that

$$
\begin{equation*}
\Omega_{2} \xrightarrow[x \rightarrow \pm \infty]{ }+\infty \tag{5.10}
\end{equation*}
$$

This implies, by the arguments of the appendix, that $g_{1}$ is positive and its inverse is normalisable.

Therefore we can construct the new potential with two added eigenvalues as

$$
\begin{align*}
V_{2} & =V_{1}-2 \frac{\mathrm{~d}^{2}}{\mathrm{~d} x^{2}} \ln g_{1} \\
& =V_{0}-2 \frac{\mathrm{~d}^{2}}{\mathrm{~d} x^{2}}\left(\ln g_{1}+\ln g_{0}\right) \\
& =V_{0}-2 \frac{\mathrm{~d}^{2}}{\mathrm{~d} x^{2}} \ln \frac{\Omega_{2}\left(g_{0}\left(\lambda_{1}, \alpha_{1}\right), g_{0}\left(\lambda_{2},-\alpha_{2}\right)\right)}{\Omega_{1}\left(g_{0}\left(\lambda_{1}, \alpha_{1}\right)\right)} \Omega_{1}\left(g_{0}\left(\lambda_{1}, \alpha_{1}\right)\right) \\
& =V_{0}-2 \frac{\mathrm{~d}^{2}}{\mathrm{~d} x^{2}} \ln \Omega_{2}\left(g_{0}\left(\lambda_{1}, \alpha_{1}\right), g_{0}\left(\lambda_{2},-\alpha_{2}\right)\right) . \tag{5.11}
\end{align*}
$$

Its ground state and the eigenfunction corresponding to $\lambda_{1}$ are

$$
\begin{align*}
\varphi_{(2), \lambda_{2}} & =c_{2,2} g_{1}^{-1}  \tag{5.12}\\
\varphi_{(2), \lambda_{1}} & =c_{2,1} A_{2}^{\dagger} g_{0}^{-1} \tag{5.13}
\end{align*}
$$

where $c_{2,2}$ and $c_{2,1}$ are normalisation constants. The normalised eigenfunctions for the eigenvalues $E_{n}>\lambda_{1}$ are given by

$$
\begin{align*}
\varphi_{(2), n} & =\left(E_{n}-\lambda_{2}\right)^{-1 / 2} A_{2}^{\dagger} \varphi_{(1), n}  \tag{5.14}\\
& =\left(E_{n}-\lambda_{2}\right)^{-1 / 2}\left(E_{n}-\lambda_{1}\right)^{-1 / 2} A_{2}^{\dagger} A_{1}^{\dagger} \varphi_{(0), n}  \tag{5.15}\\
& =Z_{2, n} B_{2}^{\dagger} \boldsymbol{\varphi}_{(0), n} . \tag{5.16}
\end{align*}
$$

In the last equality we have introduced the notation

$$
\begin{equation*}
B_{N}^{\dagger}=A_{N}^{\dagger} A_{N-1}^{\dagger} \ldots A_{1}^{\dagger} \tag{5.17}
\end{equation*}
$$

with

$$
\begin{equation*}
A_{i}^{\dagger}=-g_{i-1} \frac{\mathrm{~d}}{\mathrm{~d} x} g_{i-1}^{-1} \tag{5.18}
\end{equation*}
$$

and for the normalisation

$$
\begin{equation*}
Z_{N, n} \doteq \prod_{i=1}^{N}\left(E_{n}-\lambda_{i}\right)^{-1 / 2} \tag{5.19}
\end{equation*}
$$

Notice that the operator $B_{2}^{\dagger}$ acting on a function $f$ can be represented as

$$
\begin{equation*}
B_{2}^{\dagger} f=\frac{\Omega_{3}\left(g_{0}\left(\lambda_{1}, \alpha_{1}\right), g_{0}\left(\lambda_{2},-\alpha_{2}\right), f\right)}{\Omega_{2}\left(g_{0}\left(\lambda_{1}, \alpha_{1}\right), g_{0}\left(\lambda_{2},-\alpha_{2}\right)\right)} \tag{5.20}
\end{equation*}
$$

## 6. Adding $N$ eigenvalues

The results of the previous section are easily extended to the simultaneous addition of $N$ eigenvalues. Given $N$ arbitrary values $\lambda_{0}>\lambda_{1}>\lambda_{2}>\ldots>\lambda_{N}$ and $N$ positive constants $\alpha_{1}, \ldots, \alpha_{N}$, we construct the potential
$V_{N}=V_{0}-2 \frac{\mathrm{~d}^{2}}{\mathrm{~d} x^{2}} \ln \Omega_{N}\left(g_{0}\left(\lambda_{1},+\alpha_{1}\right), g_{0}\left(\lambda_{2},-\alpha_{2}\right), g_{0}\left(\lambda_{3},+\alpha_{3}\right), \ldots, g_{0}\left(\lambda_{N},(-1)^{N+1} \alpha_{N}\right)\right)$.

The corresponding Hamiltonian $H_{N}$ has the same spectrum as $H_{0}$ plus the eigenvalues $\lambda_{1}, \lambda_{2}, \ldots, \lambda_{N}$. Its ground state is

$$
\begin{equation*}
\varphi_{(N), \lambda_{N}}=c_{N, N} g_{N-1}^{-1} \tag{6.2}
\end{equation*}
$$

where $g_{N-1}$ is the general positive solution of $H_{N-1} g_{N-1}=\lambda_{N} g_{N-1}$, which is given by (6.6). The eigenfunctions corresponding to $\lambda_{1}, \lambda_{2}, \ldots, \lambda_{N-1}$ are

$$
\begin{equation*}
\varphi_{(N), \lambda_{i}}=c_{N, i} A_{N}^{\dagger} \ldots A_{i+1}^{\dagger} g_{i-1}^{-1} . \tag{6.3}
\end{equation*}
$$

The normalised eigenfunctions for the eigenvalues $E_{n}>\lambda_{1}$ are given by

$$
\begin{equation*}
\varphi_{(N), n}=Z_{N, n} B_{N}^{\dagger} \varphi_{(0), n} \tag{6.4}
\end{equation*}
$$

(with the notation introduced in (5.17) and (5.19)).
The operator $B_{N}^{\dagger}$ can be represented by

$$
\begin{equation*}
B_{N}^{\dagger} f=(-1)^{N} \frac{\Omega_{N+1}\left(g_{0}\left(\lambda_{1}, \alpha_{1}\right), \ldots, g_{0}\left(\lambda_{N},(-1)^{N+1} \alpha_{N}\right), f\right)}{\Omega_{N}\left(g_{0}\left(\lambda_{1}, \alpha_{1}\right), \ldots, g_{0}\left(\lambda_{N},(-1)^{N+1} \alpha_{N}\right)\right)} . \tag{6.5}
\end{equation*}
$$

This relation follows immediately from

$$
\begin{equation*}
g_{N-1}=\frac{\Omega_{N}\left(g_{0}\left(\lambda_{1}, \alpha_{1}\right), \ldots, g_{0}\left(\lambda_{N},(-1)^{N+1} \alpha_{N}\right)\right)}{\Omega_{N-1}\left(g_{0}\left(\lambda_{1}, \alpha_{1}\right), \ldots, g_{0}\left(\lambda_{N-1},(-1)^{N} \alpha_{N-1}\right)\right)} \tag{6.6}
\end{equation*}
$$

which is obtained by induction from the results of $\S 5$. (This is done, for example, in [ $5, \mathrm{p} 176]$ for the case $V_{0}=0$, but the algebra is the same in the general case.)

Therewith we obtain (6.1) by

$$
\begin{align*}
V_{N} & =V_{N-1}-2 \frac{\mathrm{~d}^{2}}{\mathrm{~d} x^{2}} \ln \frac{\Omega_{N}}{\Omega_{N-1}} \\
& =V_{N-2}-2 \frac{\mathrm{~d}^{2}}{\mathrm{~d} x^{2}} \ln \frac{\Omega_{N}}{\Omega_{N-1}} \frac{\Omega_{N-1}}{\Omega_{N-2}} \\
& \vdots  \tag{6.7}\\
& =V_{0}-2 \frac{\mathrm{~d}^{2}}{\mathrm{~d} x^{2}} \ln \Omega_{N}
\end{align*}
$$

The rest is an immediate consequence of the results of the previous sections.

## 7. Propagators

In this section we show that the time-dependent propagators $G_{N}$ and $G_{N-1}$ corresponding to the Hamiltonians $H_{N}$ and $H_{N-1}$ are related by the two following equivalent equations:

$$
\begin{align*}
G_{N}\left(x, x_{0} ; t\right)= & -g_{N-1}^{-1}\left(x_{0}\right) A_{N}^{\dagger}(x) \int_{x_{0}}^{\infty} \mathrm{d} z G_{N-1}(x, z, t) g_{N-1}(z)  \tag{7.1}\\
G_{N}\left(x, x_{0} ; t\right)= & \theta(t) \exp \left(\sigma^{-1} \lambda_{N} t\right) \varphi_{(N), \lambda_{N}}(x) \varphi_{(N), \lambda_{N}}\left(x_{0}\right) \\
& -\sigma^{-1} \theta(t) A_{N}^{\dagger}(x) A_{N}^{\dagger}\left(x_{0}\right) \int_{1}^{\infty} \mathrm{d} s G_{N-1}\left(x, x_{0}, s\right) \exp \left[\sigma^{-1} \lambda_{N}(t-s)\right] . \tag{7.2}
\end{align*}
$$

By iteration we can express $G_{N}$ in terms of the known $G_{0}$. The result is given in (7.21) and (7.22). The validity of (7.1) and (7.2) can be checked directly by insertion. They were constructed as follows.
(1) For the construction of (7.1) we forget the Hilbert space structure and just consider the differential equations. In $\S 2$ we saw that the time-dependent solutions of

$$
\begin{equation*}
\sigma \frac{\mathrm{d}}{\mathrm{~d} t} \psi_{N}(x, t)=H_{N} \psi_{N}(x, t) \tag{7.3}
\end{equation*}
$$

can be expressed in terms of solutions of

$$
\begin{equation*}
\sigma \frac{\mathrm{d}}{\mathrm{~d} t} \psi_{N-1}(x, t)=H_{N-1} \psi_{N-1}(x, t) \tag{7.4}
\end{equation*}
$$

by

$$
\begin{equation*}
\psi_{N}(x, t)=A_{N}^{+} \psi_{N-1}(x, t) \tag{7.5}
\end{equation*}
$$

Notice that in (7.5) we do not ask that the $\psi$ are in the Hilbert space.

Thus, in order to relate the propagators $G_{N}$ and $G_{N-1}$, we need to determine the initial condition $\psi_{N-1}(x, 0)$ that gives

$$
\begin{equation*}
A_{N}^{\dagger} \psi_{N-1}(x, 0)=\psi_{N}(x, 0)=\delta\left(x-x_{0}\right) \tag{7.6}
\end{equation*}
$$

The inversion of (7.6) gives

$$
\begin{align*}
\psi_{N-1}(x, 0) & =-g_{N-1}(x) \int_{-\infty}^{x} \mathrm{~d} z g_{N-1}^{-1}(z) \delta\left(z-x_{0}\right) \\
& =-g_{N-1}(x) g_{N-1}^{-1}\left(x_{0}\right)\left(1-\theta\left(x-x_{0}\right)\right) \tag{7.7}
\end{align*}
$$

( $\psi_{N-1}(x, 0)$ is clearly not in $L_{2}$.) The propagator is given by

$$
\begin{align*}
G_{N}\left(x, x_{0}, t\right) & =\theta(t) A_{N}^{\dagger}(x) \psi_{N-1}(x, t) \\
& =A_{N}^{\dagger}(x) \int_{-\infty}^{\infty} \mathrm{d} z G_{N-1}(x, z, t) \psi_{N-1}(z, 0) \tag{7.8}
\end{align*}
$$

Inserting (7.7) into (7.8) we obtain (7.1).
(2) To obtain the second expression (7.2) we work in the Hilbert space. There the general solutions of (7.3) and (7.4) are related by

$$
\begin{align*}
& \psi_{N-1}(x, t)=A_{N} \psi_{N}(x, t)  \tag{7.9}\\
& \psi_{N}(x, t)=A_{N}^{+} \psi_{N-1}(x, t)+c \varphi_{(N), \lambda_{N}}(x) \exp \left(\sigma^{-1} \lambda_{N} s\right) \tag{7.10}
\end{align*}
$$

where $c$ is a constant which is given, e.g. for the initial condition $\psi_{N}(x, 0)=\delta\left(x-x_{0}\right)$, by

$$
\begin{equation*}
c=\left\langle\psi_{N}(x, 0), \varphi_{(N), \lambda_{N}}(x)\right\rangle=\varphi_{(N), \lambda_{N}}\left(x_{0}\right) . \tag{7.11}
\end{equation*}
$$

The last term in (7.10) must be added because all the functions of the form $A_{N}^{\dagger} \phi$ are orthogonal to $\varphi_{(N), \lambda_{N}}$ :

$$
\begin{equation*}
\left\langle A_{N}^{\dagger} \phi, \varphi_{(N), \lambda_{N}}\right\rangle=\left\langle\phi, A_{N} \varphi_{(N), \lambda_{N}}\right\rangle=0 . \tag{7.12}
\end{equation*}
$$

Every function $\psi$ of $L_{2}$ (i.e. every admissible initial condition) can be expressed as $\psi=A_{N} \phi$, with the appropriate choice of $\phi$ in $L_{2}$. However, the functions of the form $\psi=A_{N}^{\dagger} \phi$ give only all the functions that are orthogonal to $\varphi_{(N), \lambda_{N}}$. Notice that in (7.5) the necessity of this supplementary term is avoided by allowing functions that are not in the Hilbert space.

Now, to obtain the second expression (7.2), instead of the inversion (7.7) we let $A_{N}$ act on (7.10) and get

$$
\begin{equation*}
\left(H_{N-1}-\lambda_{N}\right) \psi_{N-1}(x, 0)=A_{N} \psi_{N}(x, 0) . \tag{7.13}
\end{equation*}
$$

The operator ( $H_{N-1}-\lambda_{N}$ ) is invertible and we can write

$$
\begin{equation*}
\psi_{N-1}(x, 0)=R_{N-1} A_{N} \psi_{N}(x, 0) \tag{7.14}
\end{equation*}
$$

where $R_{N-1} \doteq\left(H_{N-1}-\lambda_{N}\right)^{-1}$ is the resolvent, whose integral kernel is

$$
\begin{equation*}
\tilde{R}_{N-1}(x, y)=-\sigma^{-1} \int_{0}^{\infty} \mathrm{d} s G_{N-1}(x, y, s) \exp \left(-\sigma^{-1} \lambda_{N} s\right) \tag{7.15}
\end{equation*}
$$

$G_{N}\left(x, x_{0}, t\right)$ is the solution with initial condition $\delta\left(x-x_{0}\right)$. By the preceding arguments it can be written as

$$
\begin{equation*}
G_{N}\left(x, x_{0}, t\right)=\theta(t) \varphi_{(N), \lambda_{N}}(x) \varphi_{(N), \lambda_{N}}\left(x_{0}\right) \exp \left(\sigma^{-1} \lambda_{N} s\right)+\theta(t) A_{N}^{\dagger}(x) \hat{\psi}_{N-1}\left(x, x_{0}, t\right) \tag{7.16}
\end{equation*}
$$

where $\hat{\psi}_{N-1}\left(x, x_{0}, t\right)$ is the solution of (7.4) with initial condition (according to (7.14)) $\hat{\psi}_{N-1}\left(x, x_{0}, 0\right)=R_{N-1} A_{N}(x) \delta\left(x-x_{0}\right)$. It is given by

$$
\begin{align*}
\hat{\psi}_{N-1}\left(x, x_{0}, t\right) & =\exp \left(\sigma^{-1} H_{N-1} t\right) R_{N-1} A_{N} \delta\left(x-x_{0}\right) \\
& =R_{N-1} \int_{-\infty}^{\infty} \mathrm{d} y G_{N-1}(x, y, t) A_{N} \delta\left(y-x_{0}\right) \\
& =R_{N-1} \int_{-\infty}^{\infty} \mathrm{d} y \delta\left(y-x_{0}\right) A_{N}^{\dagger}(y) G_{N-1}(x, y, t) \\
& =A_{N}^{\dagger}\left(x_{0}\right) \int_{-\infty}^{\infty} \mathrm{d} z \tilde{R}_{N-1}(x, z) G_{N-1}\left(z, x_{0}, t\right) \\
& =-\sigma^{-1} A_{N}^{\dagger}\left(x_{0}\right) \int_{t}^{\infty} \mathrm{d} s G_{N-1}\left(x, x_{0}, s\right) \exp \left[\sigma^{-1} \lambda_{N}(t-s)\right] \tag{7.17}
\end{align*}
$$

For the last equality we have used (7.15) and the identity

$$
\begin{equation*}
G_{N-1}\left(x, x_{0}, s\right)=\int_{-\infty}^{\infty} \mathrm{d} z G_{N-1}(x, z, s-t) G_{N-1}\left(z, x_{0}, t\right) \tag{7.18}
\end{equation*}
$$

This, together with (7.16), gives the result (7.2).
We remark that this result can also be obtained in a simple way in the case of a purely discrete spectrum, using the representation

$$
\begin{equation*}
G\left(x, x_{0}, t\right)=\theta(t) \sum_{n=0}^{\infty} \varphi_{n}(x) \varphi_{n}\left(x_{0}\right) \exp \left(\sigma^{-1} E_{n} t\right) \tag{7.19}
\end{equation*}
$$

and applying the transformation (2.16) to $\varphi_{n}(x)$ and $\varphi_{n}\left(x_{0}\right)$ and the identity

$$
\begin{align*}
-\sigma^{-1} \int_{t}^{\infty} \mathrm{d} s & G_{N-1}\left(x, x_{0}, s\right) \exp \left[\sigma^{-1} \lambda(t-s)\right] \\
& =\theta(t) \sum_{n=0}^{\infty} \varphi_{n}(x) \varphi_{n}\left(x_{0}\right)\left(E_{n}-\lambda\right)^{-1} \exp \left(\sigma^{-1} E_{n} t\right) \tag{7.20}
\end{align*}
$$

(3) We remark finally that after performing the iteration of (7.1) the result can be written compactly as

$$
\begin{align*}
G_{N}\left(x, x_{0} ; t\right)= & (-1)^{N} B_{N}^{\dagger}(x) \int_{x_{0}}^{\infty} \mathrm{d} z_{N} \int_{z_{N}}^{\infty} \mathrm{d} z_{N-1} \ldots \int_{z_{2}}^{\infty} \mathrm{d} z_{1} G_{0}\left(x, z_{1}, t\right) \\
& \times \prod_{i=1}^{N}\left(g_{i-1}^{-1}\left(z_{i+1}\right) g_{i-1}\left(z_{i}\right)\right) \tag{7.21}
\end{align*}
$$

with $z_{N+1} \equiv x_{0}$. The iteration of (7.2) gives (with $s_{N+1} \equiv t$ )

$$
\begin{align*}
G_{N}\left(x, x_{0} ; t\right)= & \theta(t) \sum_{i=1}^{N} \exp \left(\sigma^{-1} \lambda_{N} t\right) \varphi_{(i), \lambda_{i}}(x) \varphi_{(i), \lambda_{1}}\left(x_{0}\right) \\
& +\theta(t)(-\sigma)^{-N} B_{N}^{+}(x) B_{N}^{+}\left(x_{0}\right) \int_{i}^{\infty} \mathrm{d} s_{N} \int_{s_{N}}^{\infty} \mathrm{d} s_{N-1} \ldots \\
& \times \int_{s_{2}}^{\infty} \mathrm{d} s_{1} G_{0}\left(x, x_{0}, s_{1}\right) \exp \left(\sigma^{-1} \sum_{i=1}^{N}\left(s_{i+1}-s_{i}\right) \lambda_{i}\right) \tag{7.22}
\end{align*}
$$

## 8. Conclusion

We have thus a method to construct families of exactly solvable models depending on $2 N$ free parameters. One can use this freedom to construct potentials that have certain qualitative features that mimic models for particular physical problems.

The asymptotic behaviour is the same within a family. Thus the modifications occur in relatively localised regions. For instance, one observes that, if two eigenvalues are put close to each other, the potential develops a barrier between two wells, as one may expect from the theory of tunnel splitting.

The case $V_{0}=0$ is related to soliton solutions of the Korteweg-de Vries equation [14]. Since the number of solitons is equal to the number of eigenvalues, an appropriate choice of the parameters can give one potential well per eigenvalue.

The choice of $V_{0}$ is limited in practice by the need to know explicitly the general solution of the corresponding stationary Schrödinger equation as well as the propagator. Some examples are $V_{0}=0, V_{0}=a x^{2}$ and $V_{0}=a x^{2}+b x^{-2}$ (see $[15,16]$ ).

We remark finally that this method allows us also to construct Fokker-Planck models with prescribed properties for which the transition probability density and the eigenfunctions can be explicitly calculated. The transition probability density corresponding to the Fokker-Planck equation

$$
\begin{equation*}
\frac{\mathrm{d}}{\mathrm{~d} t} P=-\frac{\mathrm{d}}{\mathrm{~d} x}(K P)+\frac{\mathrm{d}^{2}}{\mathrm{~d} x^{2}} P \tag{8.1}
\end{equation*}
$$

with

$$
\begin{equation*}
K(x)=2 \frac{\mathrm{~d}}{\mathrm{~d} x} \ln \varphi_{(N), \lambda_{N}}(x) \tag{8.2}
\end{equation*}
$$

is given by

$$
\begin{equation*}
P\left(x, t \mid x_{0}\right)=\frac{\varphi_{(N), \lambda_{N}}(x)}{\varphi_{(N), \lambda_{N}}\left(x_{0}\right)} G_{N}\left(x, x_{0}, t\right) \exp \left(\lambda_{N} t\right) \tag{8.3}
\end{equation*}
$$

For example we consider the potential $V_{1}$ constructed by adding one eigenvalue $\lambda_{1}=-\mu^{2}$ to the free particle $\left(V_{0}=0\right)$. The general positive solution of (3.7) is

$$
\begin{equation*}
g_{0}\left(x ;-\mu^{2}, \alpha\right)=\operatorname{ch}(\mu x ; \alpha) \doteq \frac{1}{2}\left(\alpha \mathrm{e}^{\mu x}+\mathrm{e}^{-\mu x}\right) \tag{8.4}
\end{equation*}
$$

The new potential is (by (3.8)):

$$
\begin{equation*}
V_{1}=\alpha \mu / \operatorname{ch}(\mu x ; \alpha) \tag{8.5}
\end{equation*}
$$

and the normalised ground state

$$
\begin{equation*}
\varphi_{1, \lambda_{1}}(x)=(2 \mu \alpha)^{1 / 2} / \operatorname{ch}(\mu x ; \alpha) \tag{8.6}
\end{equation*}
$$

The propagator can be easily calculated using (7.1):

$$
\begin{align*}
G\left(x, x_{0} ; t\right)= & \frac{1}{(4 \pi t)^{1 / 2}} \exp \left(-\frac{\left(x-x_{0}\right)^{2}}{4 t}\right)+\frac{\mu \exp \left(\mu^{2} t\right)}{4 \operatorname{ch}\left(\mu x_{0} ; \alpha\right) \operatorname{ch}(\mu x ; \alpha)} \\
& \times\left[\operatorname{erf}\left(\frac{2 \mu t+\left(x-x_{0}\right)}{\sqrt{4 t}}\right)+\operatorname{erf}\left(\frac{2 \mu t-\left(x-x_{0}\right)}{\sqrt{4 t}}\right)\right] \tag{8.7}
\end{align*}
$$

where erf denotes the error function.

We can now consider the Fokker-Planck equation (8.1) with drift

$$
\begin{equation*}
K=-2 \mu \operatorname{th}(\mu x ; \alpha) \doteq-2 \mu \frac{\alpha \mathrm{e}^{\mu x}-\mathrm{e}^{-\mu x}}{\alpha \mathrm{e}^{\mu x}+\mathrm{e}^{-\mu x}} \tag{8.8}
\end{equation*}
$$

Its stationary state is

$$
\begin{equation*}
P_{s}=\varphi_{1, \lambda_{1}}^{2}(x)=\frac{2 \mu \alpha}{\operatorname{ch}^{2}(\mu x ; \alpha)} \tag{8.9}
\end{equation*}
$$

and the transition probability density is given by

$$
\begin{align*}
P\left(x, t \mid x_{0}\right)= & \frac{\operatorname{ch}\left(\mu x_{0} ; \alpha\right)}{\operatorname{ch}(\mu x ; \alpha)} \frac{1}{(4 \pi t)^{1 / 2}} \exp \left(-\frac{\left(x-x_{0}\right)^{2}}{4 t}\right) \exp \left(-\mu^{2} t\right) \\
& \quad+\frac{\mu \alpha}{4 \operatorname{ch}^{2}(\mu x ; \alpha)}\left[\operatorname{erf}\left(\frac{2 \mu t+\left(x-x_{0}\right)}{\sqrt{4 t}}\right)+\operatorname{erf}\left(\frac{2 \mu t-\left(x-x_{0}\right)}{\sqrt{4 t}}\right)\right] \tag{8.10}
\end{align*}
$$

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## Appendix. Asymptotic properties

In this appendix we discuss some asymptotic properties of the solutions of the Schrödinger equation

$$
\begin{equation*}
-g^{\prime \prime}+V g=\lambda g \quad \lambda<\lambda_{0} \tag{A1}
\end{equation*}
$$

where $V$ has the asymptotic behaviour (3.2) and $\lambda$ is smaller than the lowest point of the spectrum $\lambda_{0}$. We use some classical results that can be found, e.g., in [17-20].
(a) $\lambda<\lambda_{0}$ implies that $g$ has at most one zero and it is simple. The proof can be found, e.g., in [17, p 208].
(b1) There are solutions $h_{1}, h_{2}$ that have the following asymptotic behaviour at $-\infty$ :

$$
\begin{align*}
& h_{1}(x) \underset{x \rightarrow-\infty}{\text { mum }}(V-\lambda)^{-1 / 4} \exp \left(\int^{x} \mathrm{~d} y(V-\lambda)^{1 / 2}\right) \rightarrow 0  \tag{A2}\\
& h_{2}(x) \underset{x \rightarrow-\infty}{\operatorname{mum}_{m}}(V-\lambda)^{-1 / 4} \exp \left(-\int^{x} \mathrm{~d} y(V-\lambda)^{1 / 2}\right) \rightarrow \infty . \tag{A3}
\end{align*}
$$

(b2) There are solutions $h_{3}, h_{4}$ that have the following asymptotic behaviour at $+\infty$ :

$$
\begin{align*}
& h_{3}(x) \underset{x \rightarrow+\infty}{\mathrm{mm}}(V-\lambda)^{-1 / 4} \exp \left(-\int^{x} \mathrm{~d} y(V-\lambda)^{1 / 2}\right) \rightarrow 0  \tag{A4}\\
& h_{4}(x) \underset{x \rightarrow+\infty}{\mathrm{m}}(V-\lambda)^{-1 / 4} \exp \left(\int^{x} \mathrm{~d} y(V-\lambda)^{1 / 2}\right) \rightarrow \infty \tag{A5}
\end{align*}
$$

This is proven, e.g., in [19, pp 190-202].
(c) $h_{1}$ and $h_{3}$ are uniquely defined by (A2) and (A4). (This is not the case for $h_{2}$ and $h_{4}$.)

Proof. If there were another linearly independent solution with the same asymptotic behaviour, then all solutions would tend to zero in that limit, which is in contradiction with (A3) and (A5).
(d) $h_{1}$ is square integrable in an interval ( $-\infty, x_{1}$ ) for some small enough $x_{1}$ and $h_{3}$ is square integrable in an interval $\left(x_{3}, \infty\right)$ for some large enough $x_{3}$. This is an immediate consequence of the exponential decay.
(e) The following are pairs of linearly independent solutions: $\left(h_{1}, h_{2}\right),\left(h_{3}, h_{4}\right)$, and $\left(h_{1}, h_{3}\right)$.

Proof. For the first two pairs it is evident from the asymptotics. If $h_{1}$ and $h_{3}$ were linearly dependent, they would be square integrable at both limits, and thus they would be eigenfunctions, which contradicts $\lambda<\lambda_{0}$.
(f) The following behaviour follows from (b)-(e):

$$
\begin{align*}
& h_{1}(x) \underset{x \rightarrow+\infty}{\text { mam }} c_{1}(V-\lambda)^{-1 / 4} \exp \left(\int^{x} \mathrm{~d} y(V-\lambda)^{1 / 2}\right) \rightarrow \infty  \tag{A6}\\
& h_{3}(x) \underset{x \rightarrow-\infty}{\text { mum }} c_{3}(V-\lambda)^{-1 / 4} \exp \left(-\int^{x} \mathrm{~d} y(V-\lambda)^{1 / 2}\right) \rightarrow \infty \tag{A7}
\end{align*}
$$

where $c_{1}$ and $c_{3}$ are constants.
(g) $h_{1}$ and $h_{3}$ are everywhere positive.

Proof. Consider $h_{1}$. For $x \rightarrow-\infty$ it is positive by definition. Since we know that it can have at most one zero (counting multiplicity), we must exclude two situations: (i) $h_{1}$ becomes negative and approaches zero at $+\infty$; (ii) $h_{1} \rightarrow-\infty$ for $x \rightarrow+\infty$ (plus the analogue cases for $h_{3}$ ). (i) can be excluded because otherwise $h_{1}$ would be an eigenfunction. To exclude (ii) we first note that $h_{1}$ and $h_{3}$ cannot both have zeros, since otherwise one could construct a linear combination $a h_{1}+b h_{3}$ which, for $a, b>0$ and $b$ small enough, would have two zeros. Assume then, e.g., that $h_{1}$ has a zero and $h_{3}$ none. The choice of a negative $a$ and a small enough $b>0$ would lead again to a solution with two zeros. The analogue argument for the reversed situation completes the proof.
(h) Therefore, the one-parameter family of solutions

$$
\begin{equation*}
g(x, \lambda, \alpha)=\alpha h_{1}(x)+h_{3}(x) \quad \alpha>0 \tag{A8}
\end{equation*}
$$

is also everywhere positive. As we mentioned in $\S 4$, these are all the positive solutions.
(i) Properties (b) and (f) imply that the $g^{-1}$ are normalisable. This gives us the one-parameter family of positive solutions that we can use to add an eigenvalue.
(j) For $|x|$ large enough, $h_{1}$ and $h_{3}$ as well as their first derivatives are monotonic functions.

Proof. Equation (A1), together with $\lambda<\lambda_{0}$ and the positivity of $h_{1}$ and $h_{3}$, imply that there is a $y$ such that for all $|x|>|y|$ one has $h_{1}^{\prime \prime}>0$ and $h_{3}^{\prime \prime}>0$, which implies the monotonicity of the first derivatives. Finally we show that asymptotically they have a constant sign, which implies the monotonicity of $h_{1}$ and $h_{3}$ for large $|x|$. Consider $h_{1}$. There is an $x_{1}<-|y|$, at which $h_{1}^{\prime}>0$. It then decreases toward zero as $x \rightarrow-\infty$ without ever becoming negative, since then it would stay negative and $h_{1}$ would diverge. On the other side there is an $x_{2}>|y|$, at which $h_{1}^{\prime}>0$. Then it stays positive for all $x>x_{2}$. Analogous arguments apply to $h_{3}^{\prime}$.
(k) This allows us to obtain the asymptotics of the derivatives from the asymptotics of the functions [20, pp 49-53]:
$h_{1}^{\prime}(x) \xrightarrow[x \rightarrow-\infty]{\longrightarrow \longrightarrow m}\left[-\frac{1}{4} V^{\prime}(V-\lambda)^{-5 / 4}+(V-\lambda)^{1 / 4}\right] \exp \left(\int^{x} \mathrm{~d} y(V-\lambda)^{1 / 2}\right) \rightarrow 0$
$h_{3}^{\prime}(x) \underset{x \rightarrow+\infty}{m \times \sim}\left[-\frac{1}{4} V^{\prime}(V-\lambda)^{-5 / 4}-(V-\lambda)^{1 / 4}\right] \exp \left(-\int^{x} \mathrm{~d} y(V-\lambda)^{1 / 2}\right) \rightarrow 0$.

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